

Revolution of Charged Particles in a Central Field of Attraction with Emission of Radiation

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Abstract

A particle of mass nm , carrying the electronic charge $-e$, revolves in an orbit through angle ψ at distances nr from a center of force of attraction, with angular momenta nL perpendicular to the orbital plane, where n is an integer greater than 0, m the electronic mass and r_1 is the radius of the first circular orbit. The equation of motion of the n th orbit of revolution is derived, revealing that an excited particle revolves in an unclosed elliptic orbit, with emission of radiation at the frequency of revolution, before settling down, after many cycles of ψ , in a stable circular orbit. In unipolar revolution, a radiating particle settles in a circular orbit of radius nr_1 round a positively charged nucleus. In bipolar revolution, two radiating particles of the same mass nm and charges e and $-e$, settle in a circular stable orbit of radius ns_1 round a common center of mass, where s_1 is the radius of the first orbit. Discrete masses nm and angular momenta nL lead to quantization of the orbits outside Bohr's quantum mechanics. The frequency of radiation in the bipolar revolution is found to be in conformity with the Balmer-Rydberg formula for the spectral lines of radiation from the atom hydrogen gas. There is a spread in frequency of emitted radiation, the frequency in the final circle being the highest, which might explain hydrogen fine structure, as observed with a diffraction grating of high resolution. The unipolar revolution is identified with the solid or liquid state of hydrogen and bipolar revolution with the gas state.

Keywords: Angular Momentum, Circle, Ellipse, Electric Charge, Field, Force, Frequency, Mass, Orbit, Radiation, Revolution

Introduction

Revolution of a body, round a centre of force of attraction, is the most common motion in the universe. This comes with revolution of planets round the Sun, binary stars round their centre of mass, a moon round a planet or an electron round the nucleus of an atom.

The German astronomer, Johannes Kepler early in the 17th century, formulated three laws, named after him, concerning the motions of planets [1, 2]. Kepler based his laws on astronomical data painstakingly collected in 30 years of meticulous observations by the Danish astronomer Tycho Brahe, to whom he was an assistant [3]. Kepler's proposals broke with a centuries-old belief based on the Ptolemaic system advanced by the Alexandrian astronomer Ptolemy, in the 2nd century AD, and the Copernican system put forward by the Polish astronomer, Nicolaus Copernicus, in the 16th century [4, 5].

The Ptolemaic cosmology postulated a geocentric universe in which the Earth was stationary and motionless at the centre of several concentric rotating spheres, which bore (in order of distance away

from the earth) the Moon, the planets and the stars. The major premises of the Copernican system are that the Earth rotates daily on its axis and revolves yearly round the Sun and that the planets also circle the Sun. Copernicus's heliocentric theories of planetary motion had the advantage of accounting for the daily and yearly motions of the Moon, the Sun and the stars and it neatly explained the observed motions of the planets. However, the reigning dogma in the 16th century, which was that of the Roman Catholic Church, favoured the heliocentric Ptolemaic system, it abhorred the Copernican theory and it hated its adherents.

The Copernican theory had some modifications and various degrees of acceptance in the 16th and 17th centuries. The most famous adherents of Copernican systems were the Italian physicist Galileo Galilei [6] and his contemporary, the astronomer Johannes Kepler [1, 2].

By December 1609 Galileo had built a telescope of 20 times magnification, with which he discovered four of the moons circling Jupiter [7]. This showed that at least some heavenly bodies move around a centre other than the Earth. By December 1610 Galileo had observed the phases of Venus, which could be explained if Venus was sometimes nearer the Earth and sometimes farther away from the Earth, following a motion round the Sun [7].

In 1616, Copernican books were subjected to censorship by the Church [8]. Galileo was instructed to no longer hold or defend the opinion that the Earth moved. He failed to conform to the ruling of the Church and after the publication of his book titled *Dialogue on the Two Chief World Systems*, he was accused of heresy, compelled to recant his beliefs and then confined to house arrest [9]. Galileo's *Dialogue* was ordered to be burned and his ideas banned.

The ideas contained in the *Dialogue* could not be suppressed by the Roman Catholic Church. Galileo's reputation continued to grow in Italy and abroad, especially after his final work. Galileo's final and greatest work is the book titled *Discourses Concerning Two New Sciences*, published in 1638. It reviews and refines his earlier studies of motion and, in general, the principles of mechanics. The book opened a road that was to lead Sir Isaac Newton [10] to the law of universal gravitation, which linked the planetary laws discovered by astronomer Kepler with Galileo's mathematical physics [10].

Kepler [1, 2] stamped the final seal of validity on the Copernican planetary system in three laws, viz.:

- (i) The paths of the planets are closed ellipses with the sun as one focus.
- (ii) The line drawn from the sun to the planet sweeps over equal areas in equal time.
- (iii) The square of the periods of revolution (T) of the different planets are proportional to the cube of their respective mean distances (r) from the sun ($T^2 \propto r^3$)

The import of Kepler's first law is that there is no dissipation of energy in the revolution of a planet, in a closed orbit, round the Sun. Any change of kinetic energy is equal to the change of potential energy. The second law means that a planet revolves round the Sun with constant angular momentum, which is the case if there is no force perpendicular to the radius vector. From the third law ($T^2 \propto r^3$) and relationship between the centripetal force and speed ($F \propto v^2/r$), with $T = 2\pi r/v$, it is deduced that the force of attraction F on a planet is inversely proportional to the square of its distance from the Sun, as discovered by Newton in 1687 [10].

Kepler's laws played an important part in the work of the great English astronomer, mathematician and physicist, Sir Isaac Newton [10]. The laws are significant for the understanding of the orbital paths of the moon, the natural satellite of the Earth, and the paths of the artificial satellites launched from space stations.

While the orbital path of a satellite is a closed ellipse, the orbit of an electrically charged particle, round a central force of attraction, is an unclosed ellipse or a closed circle. A charged particle revolves in an unclosed orbit with emission or absorption of radiation. The energy radiated is the difference between change in kinetic energy and change in potential energy. Revolution in a circular orbit, the steady motion, is without radiation and inherently stable as there is no change in the kinetic energy or potential energy of a revolving particle.

The purpose of this paper is to derive equations of the orbit of revolution of a charged particle round a centre of force of attraction. A particle of negative charge may revolve round a relatively much heavier positively charged nucleus as a unipolar radiator or two oppositely charged particles may revolve round their centre of mass as a bipolar radiator, each particle being a pole. The equations

are used to show that the orbit of a charged particle is an unclosed (aperiodic) ellipse where it moves with constant angular momentum nL , n being an integer. The discrete masses nm , (m being the electronic mass) of revolving particles lead to quantisation of the orbits. A particle revolves with emission or absorption of radiation of discrete frequencies, in many cycles of revolution, before settling into the stable circular orbit.

Unipolar revolution under a central force

Consider a particle of charge $-e$ and mass nm at a point P and time t revolving round, anticlockwise, in an angle ψ and with velocity v in an orbit under the attraction of a positive charge Q fixed at origin O, as shown in Figure 1. Here, n is an integer greater than 0, $-e$ is the electronic charge and m the electronic mass. The particle at P executes unipolar motion under a central force at O. In unipolar revolution, a particle, as the pole of the orbit, revolves round a stationary center under a force of attraction. The orbit of motion is an unclosed (aperiodic) ellipse with emission or absorption of radiation or a closed circle without radiation.

In Figure 1, the radius vector OP makes an angle ψ with the OX axis in space. The position vector \mathbf{r} of the point P, in the direction of unit vector $\hat{\mathbf{u}}$, (radial direction) and the velocity \mathbf{v} at time t are respectively given by the equations:

$$\mathbf{r} = r\hat{\mathbf{u}} \quad (1)$$

$$\mathbf{v} = \frac{d\mathbf{r}}{dt} = \frac{dr}{dt}\hat{\mathbf{u}} + r\frac{d\hat{\mathbf{u}}}{dt} \quad (2)$$

For orbital motion in the X-Y plane of the Cartesian coordinates, $d\hat{\mathbf{u}}/dt$, the angular velocity, is given by the vector (cross) product:

$$\frac{d\hat{\mathbf{u}}}{dt} = \frac{d\psi}{dt}\mathbf{k} \times \hat{\mathbf{u}} \quad (3)$$

The angle ψ is the inclination of the radius vector OP from OX, \mathbf{k} is a constant unit vector in the Z-direction, perpendicular to the orbital plane (out of the page in Figure 1) and $(d\psi/dt)\mathbf{k}$ is the angular velocity. The velocity \mathbf{v} is:

$$\mathbf{v} = \frac{d\mathbf{r}}{dt} = \frac{dr}{dt}\hat{\mathbf{u}} + r\frac{d\psi}{dt}\mathbf{k} \times \hat{\mathbf{u}} \quad (4)$$

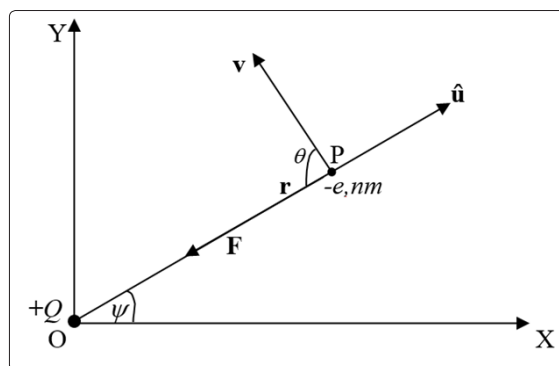


Figure 1: A particle of charge $-e$ and mass nm at a point P revolving in the X-Y plane, in angular displacement ψ , with velocity v at angle θ to force of attraction of a stationary positive charge Q at O

Noting that $(\mathbf{k} \times \hat{\mathbf{u}}) \cdot \hat{\mathbf{u}} = 0$, the velocity in the radial $\hat{\mathbf{u}}$ direction is:

$$(\mathbf{v} \cdot \hat{\mathbf{u}}) \hat{\mathbf{u}} = v_r \hat{\mathbf{u}} = \frac{dr}{dt} \hat{\mathbf{u}} \quad (5)$$

The acceleration (noting that \mathbf{k} is a constant unit vector) is obtained, a vector in two orthogonal directions, as:

$$\frac{d\mathbf{v}}{dt} = \left\{ \frac{d^2r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \right\} \hat{\mathbf{u}} + \left(2 \frac{dr}{dt} \frac{d\psi}{dt} + r \frac{d^2\psi}{dt^2} \right) \mathbf{k} \times \hat{\mathbf{u}} \quad (6)$$

The acceleration in the radial $\hat{\mathbf{u}}$ direction is:

$$\left(\frac{d\mathbf{v}}{dt} \cdot \hat{\mathbf{u}} \right) \hat{\mathbf{u}} = \left\{ \frac{d^2r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \right\} \hat{\mathbf{u}} \quad (7)$$

From equation (6), the force perpendicular to the radial direction is zero and, therefore, the acceleration of mass nm is also zero in the $\mathbf{k} \times \hat{\mathbf{u}}$ direction, so that the equation gives:

$$nm \left(2 \frac{dr}{dt} \frac{d\psi}{dt} + r \frac{d^2\psi}{dt^2} \right) \mathbf{k} \times \hat{\mathbf{u}} = \frac{nm}{r} \frac{d}{dt} \left(r^2 \frac{d\psi}{dt} \right) \mathbf{k} \times \hat{\mathbf{u}} = 0$$

This equation can be expressed in terms of angular momentum \mathbf{L} , as:

$$\begin{aligned} \frac{nm}{r} \frac{d}{dt} \left(r^2 \frac{d\psi}{dt} \right) \mathbf{k} \times \hat{\mathbf{u}} &= \frac{n}{r} \frac{d}{dt} \mathbf{L} \times \hat{\mathbf{u}} = 0 \\ n\mathbf{L} &= nmr^2 \frac{d\psi}{dt} \mathbf{k} \end{aligned} \quad (8)$$

\mathbf{L} is constant angular momentum with respect to the first orbit. Equations (7) and equation (8) will be used to derive the equation of the orbit of motion of the particle at P (Figure 1).

In Figure 1, the positively charged particle at O is supposed to be very much more massive such that it could be taken as almost stationary. In this case, we have revolution with a particle at P carrying the electronic charge $-e$ and a multiple nm of the electronic mass m , revolving in an orbit, round a stationary particle of charge $+Q$, as nucleus at O. The particle revolves in an electrostatic field \mathbf{E}_p under a force of attraction and a radiation reaction force.

The author [11, 12], invoking aberration of electric field, showed that the accelerating force \mathbf{F} on a particle of charge $-e$ and mass nm moving with velocity \mathbf{v} at a point P, in an electric field of intensity $\mathbf{E}_p = E_p \hat{\mathbf{u}}$, is:

$$\mathbf{F} = \frac{eE_p}{c} (\mathbf{c} - \mathbf{v}) = nm \left\{ \frac{d^2r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \right\} \hat{\mathbf{u}} \quad (9)$$

where the velocity of light c is at aberration angle α to \mathbf{F} and velocity \mathbf{v} is at an angle θ to \mathbf{F} .

In equation (9), taking the scalar products of accelerating force \mathbf{F} with radial vector $\hat{\mathbf{u}}$, gives:

$$\mathbf{F} \cdot \hat{\mathbf{u}} = \frac{eE_p}{c} (\mathbf{c} \cdot \hat{\mathbf{u}} - \mathbf{v} \cdot \hat{\mathbf{u}}) = nm \left\{ \frac{d^2r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \right\} \hat{\mathbf{u}} \cdot \hat{\mathbf{u}}$$

If α is a very small angle, $\mathbf{c} \cdot \hat{\mathbf{u}} = -c \cos \alpha \approx -c$ and $\mathbf{v} \cdot \hat{\mathbf{u}} = v_r = dr/dt$ (equation 5), we get:

$$-\frac{eE_p}{c} (c + v_r) = nm \left\{ \frac{d^2r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \right\} \quad (10)$$

$$-eE_p \left(1 + \frac{1}{c} \frac{dr}{dt} \right) = nm \left\{ \frac{d^2r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \right\} \quad (11)$$

where dr/dt is the speed in the radial direction, $-eE_p \hat{\mathbf{u}}$ is the electrostatic force (on a stationary particle) and $-\{eE_p (dr/dt)/c\} \hat{\mathbf{u}}$ is the radiation reaction force. The radiation reaction force, akin to a frictional force or damping force in dynamics, results in energy dissipation or radiation. Dissipation of energy or radiation always comes into play when a charged particle is accelerated or decelerated by an electric field.

Putting $E_p = Q/4\pi\epsilon_0 r^2$, in equation (11) gives the magnitude of the accelerating force as:

$$F = \frac{-eQ}{4\pi\epsilon_0 r^2} \left(1 + \frac{1}{c} \frac{dr}{dt} \right) = nm \left\{ \frac{d^2r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \right\} \quad (12)$$

$$F = \frac{-\chi}{nmr^2} \left(1 + \frac{1}{c} \frac{dr}{dt} \right) = \frac{d^2r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \quad (13)$$

where $\chi = eQ/4\pi\epsilon_0$ is a constant. Equation (13) is the mixed differential equation of motion of the particle revolving in an orbit through angle ψ and with instantaneous radius r at a time t . We need to reduce equation (13) to an equation of r as a function of one variable ψ .

In equation (13), taking the angle ψ as the variable, making the substitution $r = 1/u$ to give $dr/du = -1/u^2$ and with $(d\psi/dt) = L/mr^2$ (equation 8), we get:

$$\frac{dr}{dt} = \frac{dr}{du} \frac{d\psi}{dt} \frac{du}{d\psi} = \frac{-L}{m} \frac{du}{d\psi} \quad (14)$$

$$\frac{d^2r}{dt^2} = \frac{d}{dt} \frac{dr}{dt} = \frac{d\psi}{dt} \frac{d}{d\psi} \left(-\frac{L}{m} \frac{du}{d\psi} \right) = \frac{-L^2 u^2}{m^2} \frac{d^2u}{d\psi^2} \quad (15)$$

Substituting equations (15) and (14) into equation (13) gives:

$$\begin{aligned} -\frac{L^2 u^2}{m^2} \frac{d^2u}{d\psi^2} - \frac{L^2 u^3}{m^2} &= -\frac{\chi u^2}{nm} \left(1 - \frac{L}{mc} \frac{du}{d\psi} \right) \\ \frac{d^2u}{d\psi^2} + \frac{\chi}{ncL} \frac{du}{d\psi} + u &= \frac{m\chi}{nL^2} \end{aligned} \quad (16)$$

Equation (16) is a 2nd order differential equation with constant coefficients. A solution is $u = A \exp(x\psi)$, the transient, if the auxiliary equation, $x^2 + 2qx + 1 = 0$ and $q = \chi/2ncL$, so that:

$$x = -q \pm \sqrt{q^2 - 1} = -q \pm j\alpha$$

where α is the "rotation factor" and $\alpha^2 = 1 - q^2$, is positive. The general solution is:

$$u = \frac{1}{r} = A \exp(j\alpha - q)\psi + \frac{m\chi}{nL^2} \quad (17)$$

The appropriate solution of equation (16) is:

$$u = \frac{1}{r} = A \exp(-q\psi) \cos(\alpha\psi + \beta) + \frac{m\chi}{nL^2} \quad (18)$$

where the amplitude A , of the n th orbit, and phase angle β are obtained from the initial conditions and $m\chi/nL^2$ is the *steady state*.

Equation (18) gives the path or the n th unstable orbit of the particle (at P, in Figure 1) with O as the fixed centre of revolution or nucleus. For $q > 0 < 1$ and $\alpha < 1$, the orbit is an unclosed (aperiodic) ellipse whose major axis (line joining the points of farthest separations of the particles) rotates about an axis through the centre, perpendicular to the orbital plane. A revolving particle makes a cycle of $2\pi/\alpha$ radians as the major axis goes through $2\pi/\alpha - 2\pi$.

The exponential decay factor, $\exp(-q\psi)$, is due to energy radiation. As a result of radiation of energy, after a great number of revolutions in the angle ψ , the *transient*, that is $A \exp(-q\psi)$, decreases to zero and the radius increases to the steady state $nL^2/m\chi$, as long as q is greater than zero. This is the radius of the stable orbit when the radiator settles down from the excited state with the particle revolving in the n th stable orbit, a circle of radius $nL^2/m\chi = nr_1$, shown as *WCYD* in Figure 2. The radius r_1 is for the innermost orbit where $n = 1$.

In the stable orbit, there is only motion in a perfect circle, perpendicular to a radial electric field. No radial motion of the charged particle, no change of potential or kinetic energy and, therefore, no radiation of energy. Radiation occurs only if there is a component of velocity of a particle in the direction of an electric field.

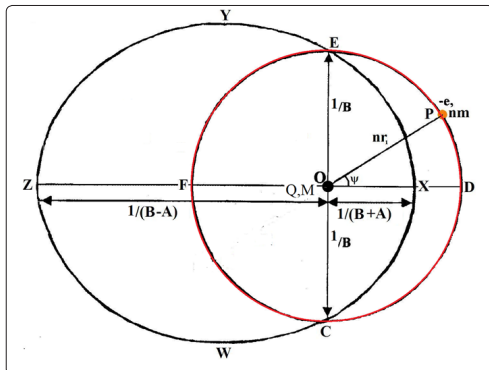


Figure 2: Free ellipse *WXYZ* of eccentricity A/B and steady orbit *CDEF* of revolution of a radiating particle, at P, in the n th circle of radius $1/B = nr_1$ with centre at O as a focus of the free ellipse

Bipolar revolution under a central force

A bipolar orbit consists of two particles of equal mass but oppositely charged, each carrying the electronic charge of magnitude e and a multiple nm of the electronic mass m , under mutual attraction, revolving round their common centre of mass, the common centre of revolution, at a point O as depicted in Figure 3. The centripetal electrostatic force of attraction F , on a charged particle, is balanced by the centrifugal force due to acceleration.

The two oppositely charged particles at P and S, in Figure 3, separated by distance $2r$, make up the two poles of the bipolar orbit, each particle being one pole in the orbit. Thus, the bipolar orbit (in contrast to the unipolar orbit) has no nucleus but an empty point as the centre of mass, the centre of revolution, located halfway between the revolving charged particles.

In Figure 3, the particle (of mass nm and charge $-e$) at point P, of position vector r , is moving with velocity v_p at an angle ψ in the electrostatic field E_p of the other particle (of mass nm and charge $+e$) at S. The particle at S of position vector $-r$, is moving with velocity v_s in the electrostatic field E_s of the particle at P. The velocities v_p and v_s are respectively:

$$v_p = \frac{dr}{dt} = \frac{dr}{dt} \hat{u} + r \frac{d\psi}{dt} \mathbf{k} \times \hat{u} \quad (19)$$

$$v_s = -\frac{dr}{dt} = -\frac{dr}{dt} \hat{u} + r \frac{d\psi}{dt} \mathbf{k} \times \hat{u} \quad (20)$$

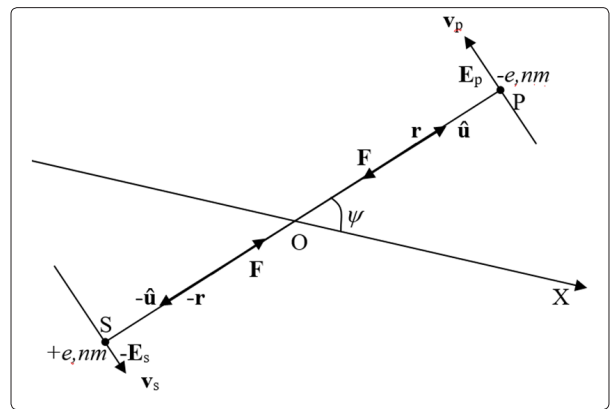


Figure 3: Two particles at P and S having electronic charges $-e$ and $+e$ and the same mass of multiple nm of electronic mass m , revolving in angle ψ , under mutual attraction, in an orbit of radius r , round the centre O.

The two particles, of equal mass, move with the same angular velocity, in a plane orbit, but with relative linear velocity in the radial direction. The relative velocity v_r of the moving particle at P with respect to the moving particle at S is:

$$v_r = v_r \hat{u} = v_p - v_s = 2 \frac{dr}{dt} \hat{u} \quad (21)$$

It is shown above that the accelerating force F , due to attraction, on a particle of charge $-e$ and mass nm revolving in an ellipse, at time t , with speed v_r in the direction of an electrostatic field $\hat{u} E_p$ of magnitude E_p (Figure 1), is given by equation (10), with $v_r = 2dr/dt$,

so that we obtain an expression similar to equation (11) as:

$$-eE_p \left(1 + \frac{2}{c} \frac{dr}{dt} \right) = nm \left\{ \frac{d^2 r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \right\} \quad (22)$$

Putting $E_p = e/16\pi\epsilon_0 r^2$, as the electric field at P, gives the force, like equation (12), as:

$$F = \frac{-e^2}{16\pi\epsilon_0 r^2} \left(1 + \frac{2}{c} \frac{dr}{dt} \right) = nm \left\{ \frac{d^2 r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \right\} \quad (23)$$

Re-arranging equation (23) gives an expression, like equation (13), as:

$$F = \frac{-\kappa}{nmr^2} \left(1 + \frac{2}{c} \frac{dr}{dt} \right) = \frac{d^2 r}{dt^2} - r \left(\frac{d\psi}{dt} \right)^2 \quad (24)$$

where $\kappa = e^2/16\pi\epsilon_0$. Equation (24) is the mixed differential equation of revolution of the particle in an orbit through angle ψ .

In equation (24), taking the angle ψ as the variable and making the substitution $r = 1/u$ and with $(d\psi/dt) = L/mr^2$ (equation 8) we get equations (14) and (15) and the first order differential equation, similar to equation (16):

$$\frac{d^2 u}{d\psi^2} + \frac{2\kappa}{ncL} \frac{du}{d\psi} + u = \frac{m\kappa}{nL^2} \quad (25)$$

If $u = (A \exp(y\psi))$ is a solution for the nth orbit, the auxiliary equation $y^2 + 2by + 1 = 0$, with $b = \kappa/ncL$. The general solution is:

$$u = \frac{1}{r} = A \exp(j\alpha - b)\psi + \frac{m\kappa}{nL^2} \quad (26)$$

The appropriate solution of equation (25) is:

$$u = \frac{1}{r} = A \exp(-b\psi) \cos(\alpha\psi + \beta) + \frac{m\kappa}{nL^2} \quad (27)$$

where the amplitude A and phase angle β are determined from the initial conditions, α is the "rotation factor" and $\alpha^2 = 1 - b^2$. The steady state is $m\kappa/nL^2$, obtained after many revolutions.

Equation (27) gives the path or the nth unstable orbit of the particles (at P or at S, in Figure 4) with O as the centre of revolution. For $b > 0 < 1$ and $\alpha < 1$, the orbit is an unclosed (aperiodic) ellipse where a particle completes one cycle of revolution in $2\pi/\alpha$ radians, with relative motion in the radial direction and with emission of radiation.

The exponential decay factor, $\exp(-b\psi)$, is due to radiation. After a great number of revolutions in the angle ψ , the transient $A \exp(-b\psi)$, decreases to zero and the radius settles at the steady state $nL^2/m\kappa$, as long b is greater than zero. This is the radius of the stable orbit when the radiating particle settles down from the excited state with the two particles revolving in the nth stable orbit, a circle of radius $nL^2/m\kappa = nr_1$, shown as WCYD in Figure 4. In the stable orbit there is only revolution in a perfect circle and no radiation.

Free ellipse and stable orbit of revolution of a radiating particle

Equation (27), of the bipolar orbit, with the phase angle β being 0, may be written as:

$$\frac{1}{r} = A \exp(-b\psi) \cos(\alpha\psi) + \frac{1}{nr_1}$$

where $r_1 = L_2/m\kappa$ is the radius of the first orbit. If b is negligible, $\alpha \approx 1$, the equation becomes:

$$\frac{1}{r} = A \cos \psi + \frac{1}{nr_1} = \frac{1}{nr_1} (1 + Ar_1 \cos \psi) \quad (28)$$

The orbit, shown as WXYZ in Figure 4, an ellipse of eccentricity $\eta = Ar_1 = A/B$, is the **free ellipse**. This is a hypothetical orbit that the particle would have taken if $b = 0$.

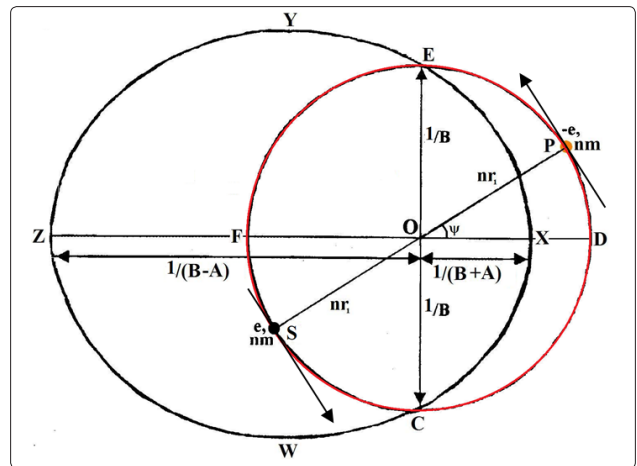


Figure 4: Free ellipse WXYZ of eccentricity A/B and steady orbit CDEF of revolution of a radiating particle, at P or S, in the nth circle of radius $n/B = nr_1$ with centre at O as a focus of the free ellipse

Revolution of a radiating particle is in an unclosed (aperiodic) ellipse, with a decreasing period (increasing frequency). After a great number of revolutions ($\psi \rightarrow \infty$), the nth orbit reduces to a circle, the **stable orbit** of radius nr_1 , shown as CDEF in Figure 4. The frequencies of revolution of a radiating particle, in the unstable orbits, are very nearly equal to that of revolution in the nth stable orbit. So, radiation from a particle, in a bipolar orbit or unipolar orbit, is a narrow band of frequencies, very nearly equal to the frequency of revolution in the nth stable orbit. This leads to a spread of frequencies of revolution, as discussed in section 5 below.

Energy radiated by a charged particle in bipolar revolution

The magnitude of accelerating force on a particle of mass nm and charge $-e$ revolving at time t and at a point distance r from the centre of a force of attraction due to an electric field of magnitude E_p of an electric charge, is given by equation (23). The radiation force is:

$$-\mathbf{R}_f = eE_p \left(\frac{2}{c} \frac{dr}{dt} \right) \hat{\mathbf{u}} = \frac{2\kappa}{c} \frac{1}{r^2} \left(\frac{dr}{dt} \right) \hat{\mathbf{u}}$$

where $\kappa = e^2/16\pi\epsilon_0$. Energy radiated is obtained by integrating the radiation force with respect to displacement (dr) in one cycle, s to $(s+1)$, through $2\pi/\alpha$ radians, of the nth orbit, to give:

$$s_r = \frac{2\kappa}{c} \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \frac{1}{r^2} \left(\frac{dr}{dt} \right) (dr)$$

Substituting for (dr/dt) from equation (14) and with $(dr) = -1/u^2 (du)$, we obtain:

$$s_r = \frac{2\kappa L}{mc} \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \left(\frac{du}{d\psi} \right)^2 (d\psi) \quad (29)$$

Substituting for $(du/d\psi)$ from equation (26), gives the integral in complex form. The energy radiated in the $(s+1)$ th cycle of the n th orbit, is given by the [Real Part] of the integral:

$$s_r = \frac{2\kappa L}{mc} \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} A^2 \exp 2\{(j\alpha - b)\psi\} (j\alpha - b)^2 (d\psi) \quad (30)$$

$$s_r = \frac{2\kappa LA^2}{mc} \left[\frac{\exp 2\{(j\alpha - b)\psi\} (j\alpha - b)^2}{2(j\alpha - b)} \right]_{\psi=\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}}$$

$$s_r = \frac{\kappa LA^2}{mc} \left[\frac{\exp(-2b\psi) \{ \cos(2\alpha\psi) + j \sin(2\alpha\psi) \} (j\alpha - b)}{1} \right]_{\psi=\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}}$$

The [Real Part] is obtained as:

$$s_r = \frac{-\kappa LA^2}{mc} \left[\frac{\exp(-2b\psi) \{ b \cos(2\alpha\psi) + \alpha \sin(2\alpha\psi) \}}{1} \right]_{\psi=\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}}$$

$$s_r = \frac{\kappa LA^2 b}{mc} \exp\left(\frac{-4\pi bs}{\alpha}\right) \left\{ 1 - \exp\left(\frac{-4\pi b}{\alpha}\right) \right\} \quad (31)$$

In the final cycle $(s \rightarrow \infty)$, the energy radiated is 0. The total energy radiated E_r , in the n th orbit, after many revolutions ending in the n th stable circle, is the sum of geometric series:

$$E_r = \sum_{s=0}^{\infty} s_r = \frac{\kappa LA^2 b}{mc} = \frac{A^2 \kappa^2}{mc^2 n} = \frac{R^2 \kappa^2}{mc^2 n^3} \quad (32)$$

where $b = \kappa/nL$ for the n th orbit, A is the amplitude for the n th orbit and R for the first orbit.

The total energy radiated by the atom is obtained by summing E_r for N orbits.

Period and Spread of frequency of oscillation of a radiator

Equation (27) gives the bipolar orbit of a charged particle of mass nm revolving through angle ψ , in the n th orbit, with constant angular momentum nL and with phase angle $\beta = 0$, as

$$\frac{1}{r} = \frac{A}{n} \exp(-b\psi) \cos(\alpha\psi) + \frac{m\kappa}{nL^2} \quad (33)$$

$$\frac{1}{r} = \frac{m\kappa}{nL^2} \left\{ 1 + \frac{AL^2}{m\kappa} \exp(-b\psi) \cos(\alpha\psi) \right\}$$

$$\frac{1}{r} = \frac{B}{n} \left\{ 1 + \frac{A}{B} \exp(-b\psi) \cos(\alpha\psi) \right\}$$

$$\frac{1}{r} = \frac{B}{n} \{ 1 + \eta \exp(-b\psi) \cos(\alpha\psi) \} \quad (34)$$

where $B = m\kappa/L^2$. This is the equation of an unclosed (aperiodic) ellipse, in the polar coordinates, with $\eta = A/B$ as the eccentricity of the free ellipse. Equation (34) gives r as:

$$r = \frac{n}{B} \{ 1 + \eta \exp(-b\psi) \cos(\alpha\psi) \}^{-1} \quad (35)$$

The angular momentum nL of the particle, of mass nm , gives:

$$dt = \frac{mr^2}{L} (d\psi)$$

$$dt = \frac{mn^2}{LB^2} \{ 1 + \eta \exp(-b\psi) \cos(\alpha\psi) \}^{-2} (d\psi) \quad (36)$$

The period of revolution $T_{(s+1)}$ in the $(s+1)$ th cycle $(s = 0, 1, 2, 3, \dots, \infty)$ of the n th orbit $(n = 1, 2, 3, \dots, N_h)$, is obtained by integrating equation (36) for y through an angle $2\pi/\alpha$ radians:

$$T_{(s+1)} = \frac{mn^2}{LB^2} \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \{ 1 + \eta \exp(-b\psi) \cos(\alpha\psi) \}^{-2} (d\psi) \quad (37)$$

Expanding the integrand, equation (37), into an infinite series, by the binomial theorem, gives:

$$\begin{aligned} & \{ 1 + \eta \exp(-b\psi) \cos(\alpha\psi) \}^{-2} \\ &= \sum_{p=0}^{\infty} (-p)^p (1+p) \eta^p \exp(-pb\psi) \cos^p(\alpha\psi) \end{aligned}$$

where p is a positive integer, $0 - \infty$. Equation (39) then becomes:

$$T_{(s+1)} = \frac{mn^2}{LB^2} \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \sum_{p=0}^{\infty} (-p)^p (1+p) \eta^p \exp(-pb\psi) \cos^p(\alpha\psi) (d\psi) \quad (38)$$

$$T_{(s+1)} = \frac{mn^2}{LB^2} \sum_{p=0}^{\infty} Q_p$$

$$Q_p = (-1)^p (1+p) \eta^p \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \exp(-pb\psi) \cos^p(\alpha\psi) (d\psi) \quad (39)$$

Expressing $\cos p(\alpha y)$ as a sum of cosines of multiples of (αy) , let us take the first five terms of Q_p .

$$Q_0 = \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} (d\psi) = \frac{2\pi}{\alpha} \quad (40)$$

$$Q_1 = -2\eta \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \exp(-b\psi) \cos(\alpha\psi) (d\psi)$$

Putting $\exp(-b\psi) \cos(\alpha\psi) = [\text{Real Part}]$ of $\{\exp(j\alpha - b)\psi\}$, the integral, Q_1 , is obtained as:

$$Q_1 = -2\eta \left[\frac{\exp(-b\psi) \{-b \cos \alpha\psi + \alpha \sin(\alpha\psi)\}}{a^2 + b^2} \right]_{\psi = \frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}}$$

Noting that $a^2 + b^2 = 1$, we get:

$$Q_1 = -2\eta b \exp\left(\frac{-2\pi bs}{\alpha}\right) \left\{ 1 - \exp\left(\frac{-2\pi b}{\alpha}\right) \right\} \quad (41)$$

$$Q_2 = -3\eta^2 \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \exp(-2b\psi) \cos^2(\alpha\psi) (d\psi)$$

$$Q_2 = \frac{-3\eta^2}{2} \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \exp(-2b\psi) \{1 + \cos(2\alpha\psi)\} (d\psi)$$

$$Q_2 = \frac{3\eta^2}{2} \left(\frac{1}{2b} + \frac{b}{2}\right) \exp\left(\frac{-4\pi bs}{\alpha}\right) \left\{ 1 - \exp\left(\frac{-4\pi b}{\alpha}\right) \right\} \quad (42)$$

$$Q_3 = -4\eta^3 \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \exp(-3b\psi) \cos^3(\alpha\psi) (d\psi)$$

$$Q_3 = -\eta^3 \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \exp(-3b\psi) \{3 \cos(\alpha\psi) + \cos(3\alpha\psi)\} (d\psi)$$

$$Q_3 = \eta^3 \left(\frac{9b}{\alpha^2 + 9b^2} + \frac{b}{3}\right) \exp\left(\frac{-6\pi bs}{\alpha}\right) \left\{ 1 - \exp\left(\frac{-6\pi b}{\alpha}\right) \right\} \quad (43)$$

$$Q_3 = 5\eta^4 \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \exp(-4b\psi) \cos^4(\alpha\psi) (d\psi)$$

Expressing $\cos^4(\alpha\psi)$ in terms of $\cos(2\alpha\psi)$ and $\cos(4\alpha\psi)$, gives:

$$Q_4 = \frac{5\eta^4}{8} \int_{\frac{2\pi s}{\alpha}}^{\frac{2\pi(s+1)}{\alpha}} \exp(-4b\psi) \{3 + \cos(4\alpha\psi) + 4 \cos(2\alpha\psi)\} (d\psi)$$

Q_4 is obtained as:

$$Q_4 = \frac{5\eta^4}{8} \left(\frac{3}{4b} + \frac{b}{4} + \frac{4b}{\alpha^2 + 4b^2}\right) \exp\left(\frac{-8\pi bs}{\alpha}\right) \left\{ 1 - \exp\left(\frac{-8\pi b}{\alpha}\right) \right\} \quad (44)$$

Note that as $s \rightarrow \infty$ or $b = 0$, only Q_0 remains, since $Q_1 = Q_2 = Q_3 = \dots = Q_p = 0$. Where $b \neq 0$, the period of the $(s + 1)$ th cycle in the n th orbit of revolution, is obtained as the sum:

$$T_{(s+1)} = \frac{mn^2}{LB^2} \sum_{p=0}^{\infty} Q_p = \frac{mn^2}{LB^2} (Q_0 + Q_1 + Q_2 + Q_3 + \dots + Q_p + \dots + Q_{\infty}) \quad (45)$$

Since the eccentricity η is small and b is smaller, neglecting powers of η greater than 2 and powers of b greater than 1, we obtain $Q_1 \approx Q_3 \approx Q_5 \approx \dots \approx Q_{2p+1} \approx 0$. An approximate expression for the period of revolution in the first cycle (with $s = 0$), in the n th orbit, is obtained as the sum:

$$T_1 = \frac{mn^2}{LB^2} \sum_{p=0}^{\infty} Q_p \approx Q_0 + Q_2 + \dots + Q_{2p} + \dots + Q_{\infty} \approx Q_0 + Q_2 \approx \frac{mn^2}{LB^2} \left[\frac{2\pi}{\alpha} + \frac{3\eta^2}{2} \left(\frac{1}{2b} + \frac{b}{2}\right) \left\{ 1 - \exp\left(\frac{-4\pi b}{\alpha}\right) \right\} \right] \quad (46)$$

With $\alpha^2 = 1 - b^2$, we obtain $\alpha \approx 1$ and:

$$T_1 \approx \frac{mn^2}{LB^2} \left(\frac{2\pi}{\alpha} + 3\eta^2 \frac{2\pi}{\alpha} \right) \approx \frac{mn^2}{LB^2} \left(2\pi + \frac{3\eta^2}{2} 2\pi \right) \quad (47)$$

Equation (47) gives the time taken, in the first cycle, for the particle, in the n th orbit, to go through $2\pi/\alpha - \pi$ radians, due to the rotation of the major axis that goes through $(2\pi/\alpha - \pi)$ radians. After a great number of cycles ($s \rightarrow \infty$), the period of revolution, through 2π radians, in the steady circle of the n th orbit, is:

$$T_n = \frac{2\pi mn^2}{LB^2} \quad (48)$$

T_n is the period of revolution of the particle in the steady orbit, a circle of radius $r_n = n/B = 16\rho n \epsilon L^2/me$ with speed $v_n = BL/nm = e^2/16\pi n \epsilon_0 L$ and frequency $f_n = LB^2/2\pi mn^2 = me^4/2\pi L^{3b^n} (16\pi n z_0)^2$.

The period of the first cycle T_1 , (equation (47)), is the highest while that in the steady orbit T_n , (equation (48)) is the least (highest frequency). The wavelength, λ_1 of radiation, during the 1st cycle of the n th orbit, is:

$$\lambda_1 = cT_1 = \frac{2\pi mn^2 c}{LB^2} \left(1 + \frac{3\eta^2}{2} \right) \quad (49)$$

where c is the speed of light in a vacuum. The separation, splitting or increment of the wavelengths $\Delta\lambda$ is:

$$\Delta\lambda = \frac{3\eta^2 \pi m n^2 c}{LB^2} \quad (50)$$

The ratio of the separation of wavelengths and the wavelength, with respect to revolution in the n th stable orbit, the same as the magnitude of separation of frequencies, is:

$$\frac{\Delta\lambda}{\lambda_n} = -\frac{\Delta f}{f_n} = \frac{3\eta^2}{2} \quad (51)$$

This ratio is related to the splitting, spread or “fine structure” of a spectral line due to frequency of radiation from the hydrogen atom. The radiation is not of a precise frequency but has a spread around the frequency of revolution of a particle in the n th stable orbit.

Results and Discussion

1. A charged particle revolves round a centre of attraction, in an unclosed (aperiodic) elliptic orbit, with emission of radiation at the frequency of revolution, before settling into a stable circular orbit. This explains the source of atomic radiation and stability of atoms.
2. The unipolar revolution of an electron round a nucleus, as discussed in section 2, leads to the development of unipolar model or nuclear model of the hydrogen atom. Similarly, bipolar motion of two oppositely charged particles of the same mass round a centre of revolution, as discussed under section 3, leads to the development of bipolar model or non-nuclear model of the hydrogen atom.
3. It is shown by the author that the mass of a particle is independent of its speed [11]. Therefore, in the treatment of the motion of electrons round a centre of revolution, in the unipolar or bipolar models of the hydrogen atom, relativistic effects were not taken into consideration. Neither was the spin of a revolving particle regarded in the emission of radiation from the hydrogen atom.
4. The narrow spread of frequencies, with respect to revolution in the n th orbit, as given by equation (51), may explain the “fine structure” of the spectral lines of radiation from the hydrogen atom, without considering relativistic effects, electron spin or quantum mechanics [12, 13].

Conclusion

The paper concludes that revolution of a charged particle in a circular orbit, round a centre of force of attraction, is inherently stable. Radiation takes place only if the particle is excited by being dislodged from the stable circular orbit. An excited particle revolves in an aperiodic elliptic orbit, emitting radiation before reverting into the stable circular orbit. This explains the source of radiation from atomic particles, outside quantum mechanics.

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